Photon Sieve Telescopes

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Photon Sieve

- Essentially a Fresnel Zone Plate with rings broken up into individual holes
- In simplest version holes are same diameter (d) as ring width (w)

$$r_n^2 = 2nf\lambda + n^2\lambda^2$$

 Can be randomly or regularly distributed with angle

$$w = \frac{\lambda f}{2r_n}$$

 Can have any density (fill) in each zone as desired

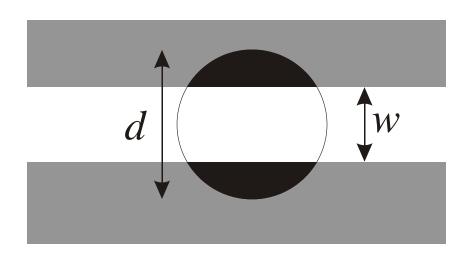


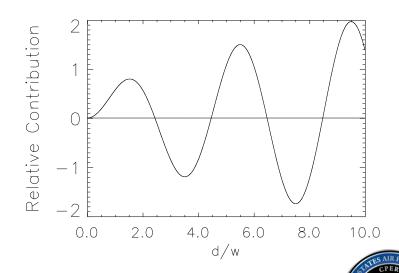
Hole size

Can make hole size (d) > underlying zone (w):

$$F \propto \frac{d}{w} J_1 \left(\frac{\pi d}{2w} \right)$$

 Still get positive contribution so long as overlap with bright zone is greater than overlap with dark zone

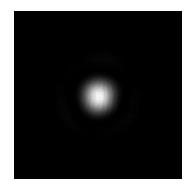


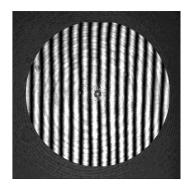


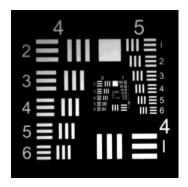
Summary of past work

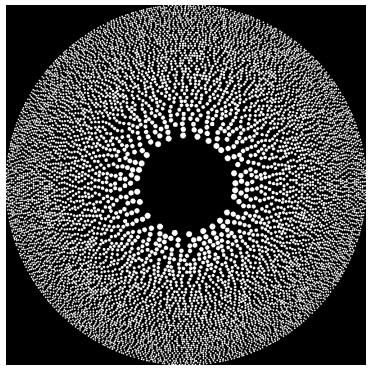
- 4" tests using chrome-coated quartz sieves
 - 5 million holes, 20-330μm in diameter
 - 0.02λ RMS wavefront, 0.98 Strehl













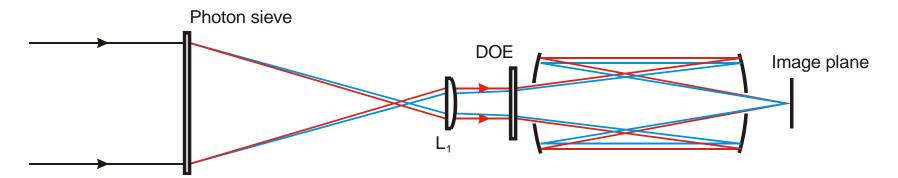
Broadband operation

• Photon sieve is narrowband due to dispersion:

$$\Delta \lambda = \frac{2\lambda f^2}{D^2}$$

Correct with secondary DOE (hologram)

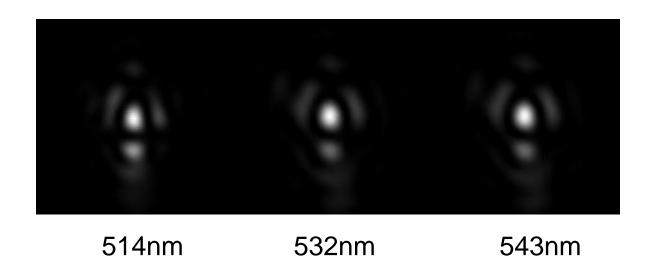






Broadband telescope

- HOE created in lab: D = 40mm, f = -158mm
- Demonstrated perfect imaging over 40nm bandwidth



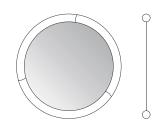


Membrane photon sieves

Quartz not an option for space applications

• Instead, want to pattern sieve on membrane

- · Can be rolled, folded into compact package
- Deployed to larger aperture in space
- Easier to pull flat than form 3D in space
 - Use inflation or tensioning





Thermal Distortion

- In-plane stretching or shrinkage of the substrate will move the holes locations off the zones
- Due to substrate itself or the support structure
 - Not simply a matter of finding a zero CTE polymer
- For a given $CTE(\alpha)$ there will be a shift in the position of the outer ring of Δr :

$$\Delta r = \alpha . r . \Delta T < w/10$$

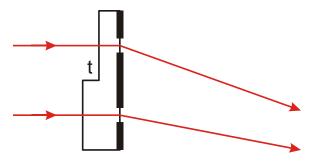
Typically in the order of 10⁻⁶ °C⁻¹



Thickness Variation

- Sieve surface may be flat but the phase of the input wavefront is affected by substrate thickness variations
 - Problem for transmissive photon sieves
- OPD between any two zones depends on material thickness, t and refractive index, n:

$$OPD = t(n-1)$$



• Max thickness error for n = 1.45:

 $\lambda/10$: t < 0.22 λ (0.11 microns)

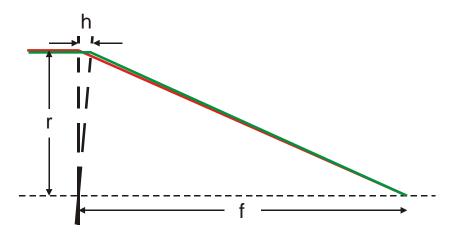


Mechanical Deformation

- Surface deformations mean diffracted light is no longer perfectly in phase from one hole to another
- A deviation of height h will change the path length of a ray from that point according to:

$$h = 2\phi (f/r)^2$$

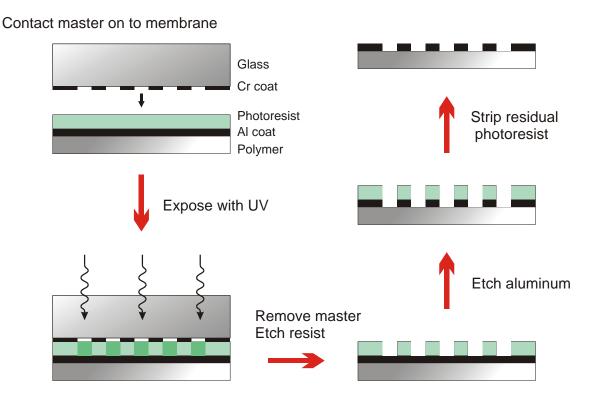
- For D=1m (f/5), $\phi = \lambda/10$:
 - Conventional mirror: $h = \lambda/20$
 - Photon sieve: $h = 20\lambda$





Polyimide film

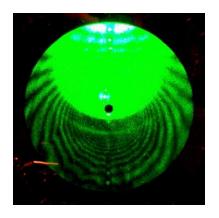
- 25µm thick polyimide with high thickness uniformity
- Coated with Al and photoresist for contact printing:

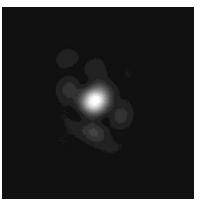


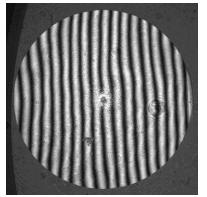


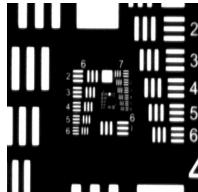
Polyimide film

- Diffraction limited performance
- 0.056λ RMS wavefront error, 0.88 Strehl
 - Even with less than perfect surface flatness evident
- Efficiency of this antihole design was just 0.35%





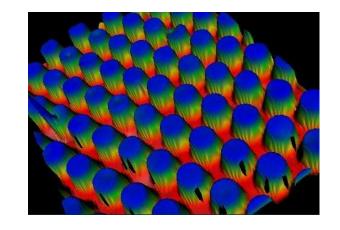






Intensity vs Phase

- Transmission photon sieves are binary intensity DOEs with limited diffraction efficiency
- Created photon sieve with optimum 50% fill
 - 3.8 million holes ranging in size from 8-395μm
- Al-coat CP1 films to convert to binary phase DOE
 - Al coating had to be $\lambda/4$ thick
 - 133nm for 532nm light

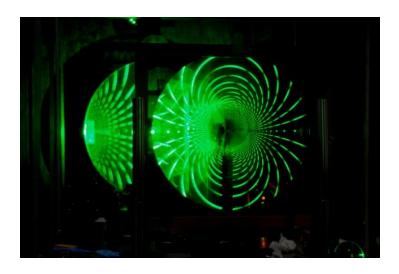


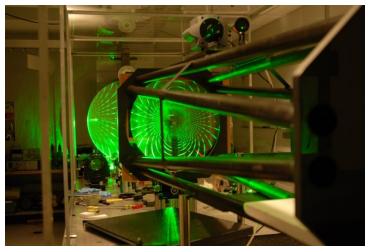
Efficiency improved from 3.5% to 10%



Other Work

- Have created 0.56m version
 - f=3m, 780 million holes!



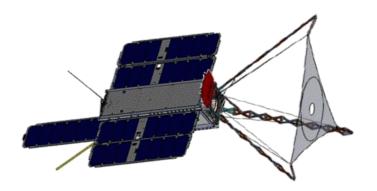


- Formed the basis of MOIRE program:
 - Goal: 20m membrane in GEO (1000kg launch mass)
 - Phase I: Meter-scale demo



FalconSAT-7

- FalconSAT program: cadets design, build, launch and operate small satellite
 - Astro Eng., Physics, Elec Eng., Management, Mech Eng.
- FalconSAT-7: Deploy membrane photon sieve to observe the Sun
 - 3U CubeSat (30cm x 10cm x 10cm)
 - 0.2m membrane photon sieve





FalconSAT-7 Team

 USAFA – Management, systems integration, science, optics, and electronics



- NRO Colony II CubeSAT, launch & networked ground stations
- NASA / Goddard Science





- MMA Design LLC Deployment system design
- AFIT CubeSAT bus and mission modeling



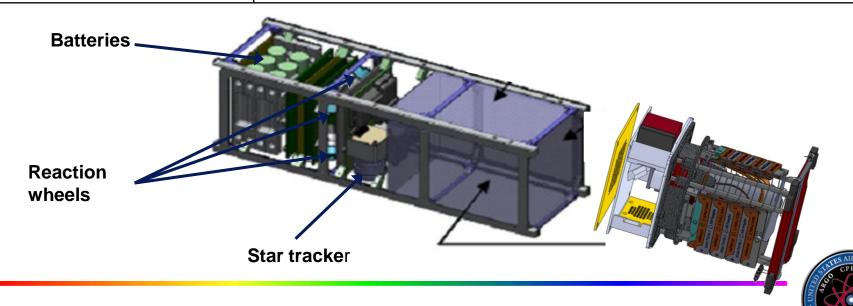


AFRL/RVSV – Membrane analysis



Colony II Spacecraft

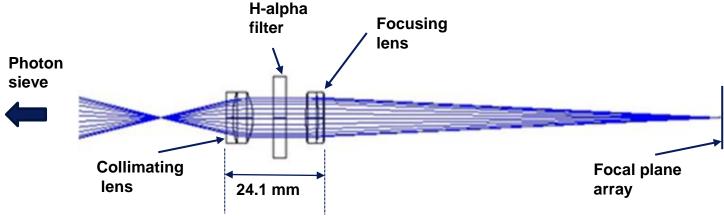
	Capability		
Launch Vehicle Interface	Adheres to CubeSat P-POD requirements		
Attitude Pointing	0.37° at all times (2σ)		
Position Knowledge	±5 km in x, y, and z		
Data Rate	50 kbps (between 440-460 MHz and 900-928 MHz		
Payload Orbit Average Power	17.5 W (50% duty cycle)		
Peak Power	Peak Power 70 W (for 2 minutes)		
Thermal Control	Temperature range is -33°C to +71°C		



Photon Sieve

- Binary phase photon sieve
 - Master/contact process or direct-write
- 2.5 billion pinholes (2-277μm diam.)
- D = 200mm, f = 400mm, λ = 656.3nm
- 50% fill factor, η ~30%

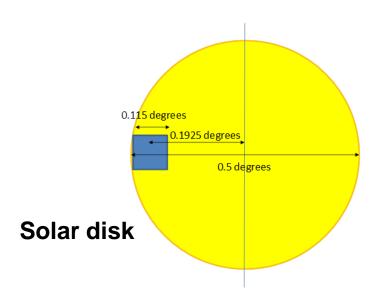


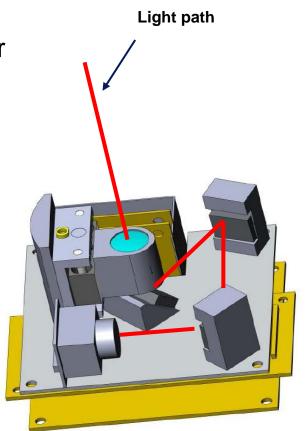




Peregrine Optical System

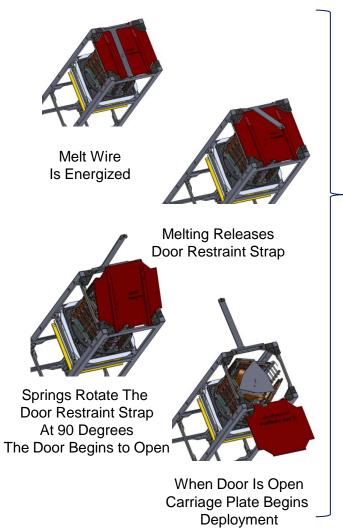
- 2 secondary lenses & H-alpha filter
 - Kinematic adjustment for focus/decenter
- 4 μrad resolution (600 km at Sun)
- ~0.1° FOV, 1Å spectral bandwidth
 - Depends on final choice of camera

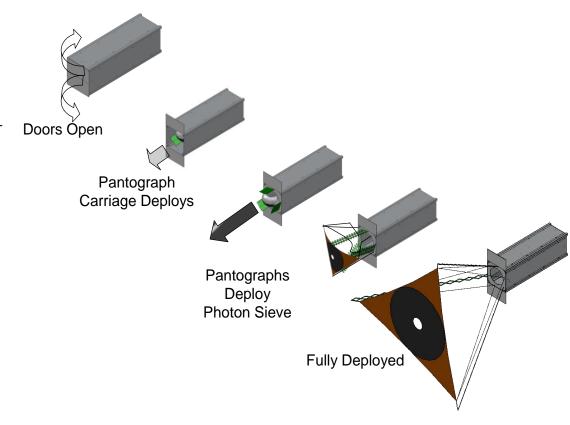






Peregrine Deployment Sequence







Project Schedule

Reviews

- Oct 2010 System Requirements Review
- Dec 2010 Conceptual Design Review
- Dec 2011 Preliminary Design Review
- May 2012 Critical Design Review

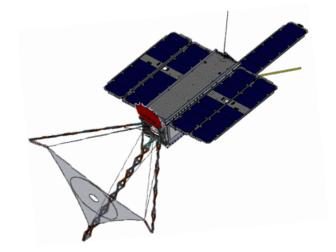
Engineering Model

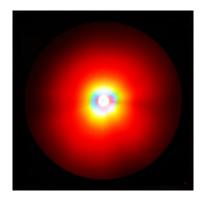
- Aug 2011 Engineering Bus Arrives
- Jun 2011 Payload Lab Prototype
- Qual Model
- May 2012 Payload Qual

Flight Model

- Oct 2011 Flight Bus Arrives
- Dec 2012 Payload Flight Model Finished
- Aug 2013 Payload Integration & Test Complete

Ready for Launch Sep 2013





Optical test pattern of photon sieve master

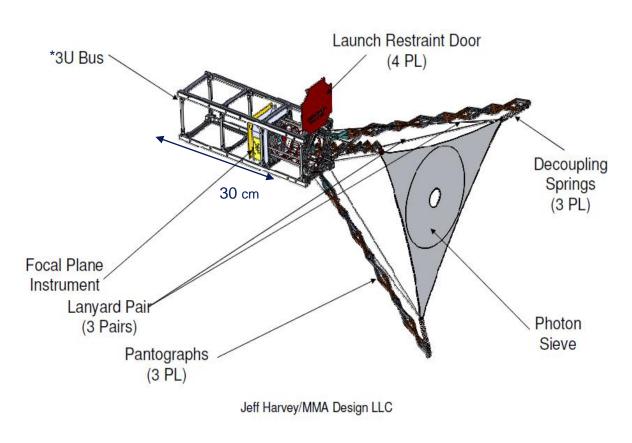


Questions?

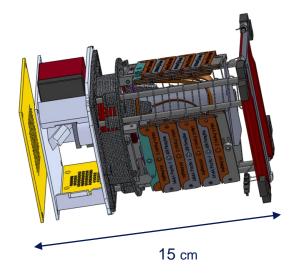




FalconSAT-7 System



Peregrine Payload



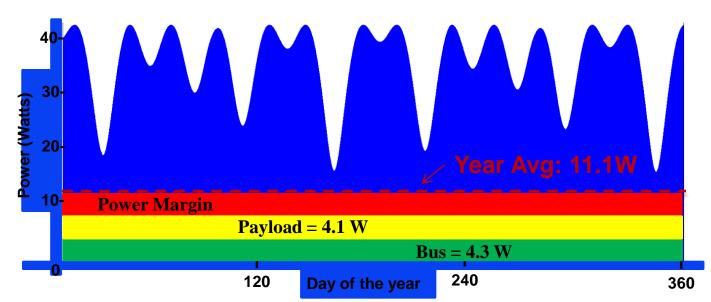
* Avionics not shown



Power

- Power analysis done by AFIT
 - Generated using STK solar panel module (28% eff. cells at 28°C)
 - Sun Beta Angle of 30 degrees

Bus Power Supply Over 1 Year



- Conclusions
 - AFIT analysis verifies Boeing power estimations
 - Removal of one solar panel will not jeopardize mission
 - Payload requirement: Peak: 6.7 W Average: 4.1 W

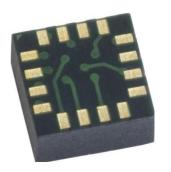
Power margin: Peak: 63.3 W Average: 2.7 W



ADCS

- AFIT review
 - Assumes baseline Colony-II
 - 700km orbit altitude
- Conclusions
 - Colony II baseline configuration meets FS-7 ADCS requirements
 - Future analysis:
 - Performance at different altitudes
 - Deployed optics effects on ADCS





3 Axis Magnetic Sensor by Honeywell

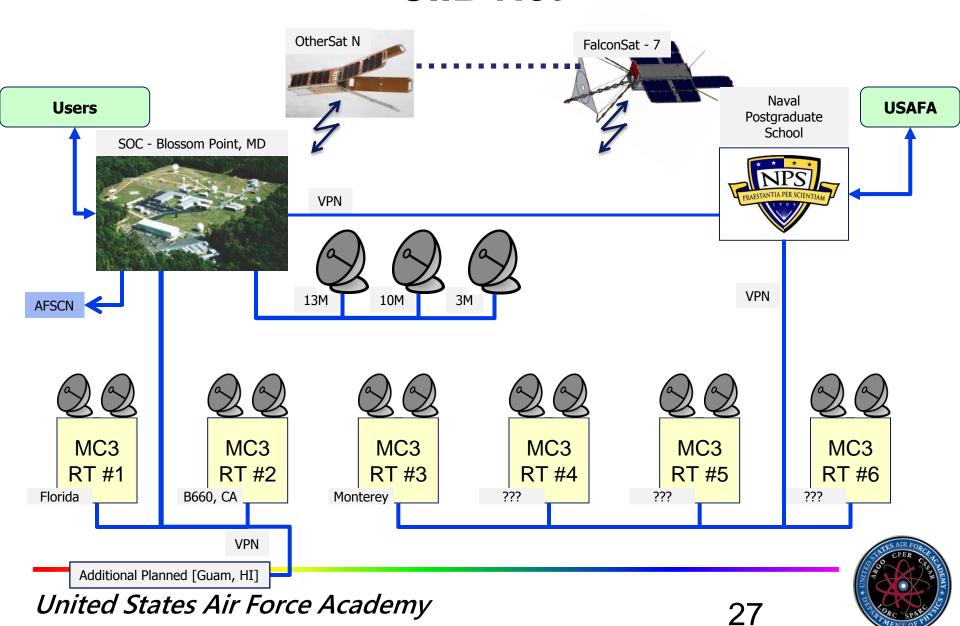


Buy CIIB Ground Station

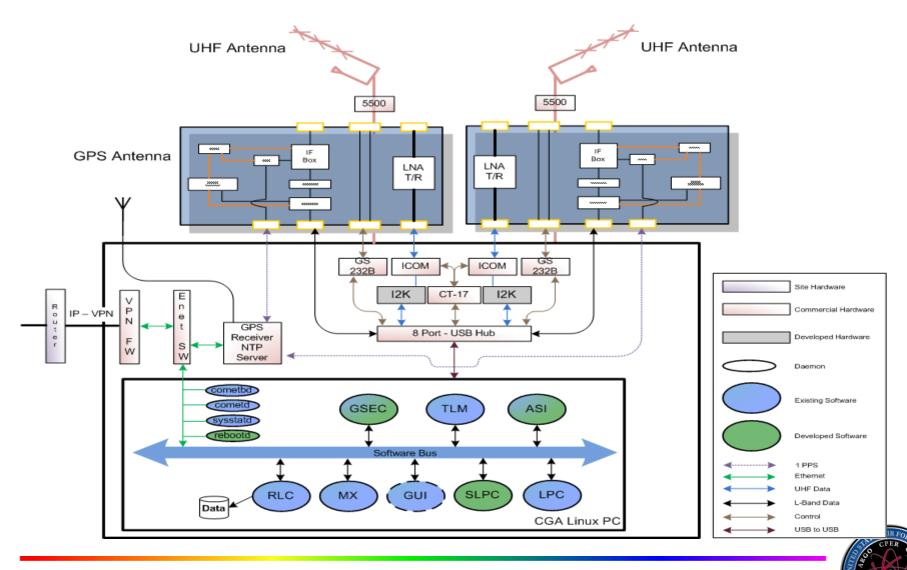
- Prebuilt CIIB GS from NRL.
- Build into current USAFA SatComm architecture
- Advantages: Additional GS at USAFA, singular source of control and decision making over comm. Time with our satellite, access to CIIB network, other ground stations.
- Disadvantages: Cost (80K-100K), already don't have man hours available, adding to complexity of ground architecture, increasing risk.



CIIB Net

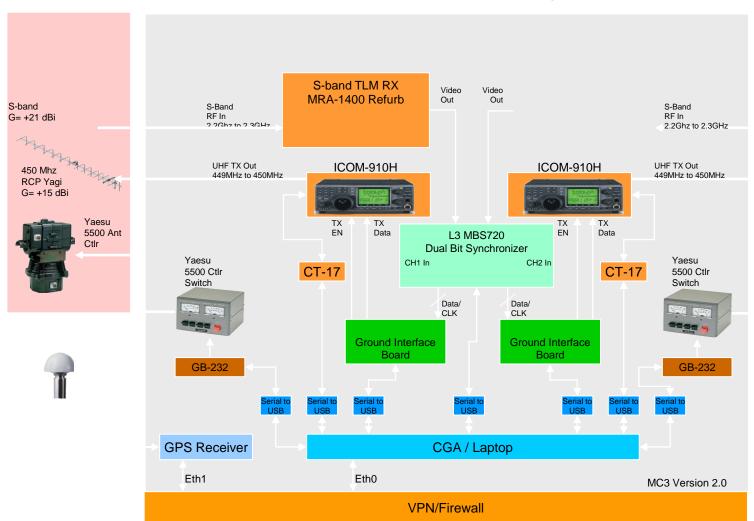


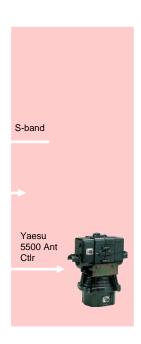
CIIB Communications System



CIIB GS Architecture

MC3 Version 2.0 Baseline Configuration





External Network
Connection To
Blossom Point AND

Communication

- Colony II Specifications
 - Uplink 450 MHz at 9.6 kbps; Downlink 915 MHz at 57.6 kbps; BER: 1E-5
 - Up to 2 GB storage on bus (shared between payload/bus)
- Payload Requirements

	MBytes/Image		Images/Day		MBytes/Day	
	Threshold	Objective	Threshold	Objective	Threshold	Objective
Sun Imaging	0.346	2	1	10	0.346	20
Earth Imaging	0.346	2	0	1	0	2

Payload has additional 2 GB storage (TBR)

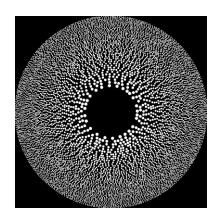
- RS-422 serial connection to bus
- Communication Analysis
 - Orbit: 450 km i=98°
 - Assume USAFA ground station
 - Sufficient link margin for up/downlink: ~10 dB above required 9.6 dB
 - ~53 sec to download one picture (10% overhead)

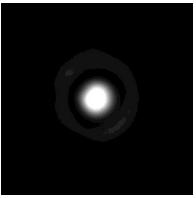


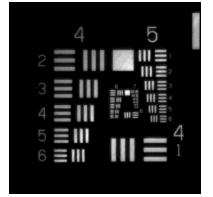


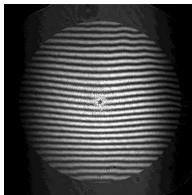
Metal film

- Made on 12.5μm thick electroformed nickel
 - D= 100mm, f=1m, 5 million antiholes
- Diffraction limited imaging but films too delicate for practical purposes
 - Thicker film not possible with 19μm holes









Diazo film

- Photosensitive polymer for lithography
 - Contact copying of quartz master
- Performance was less than perfect
 - Thick photosensitive layer resulted in blurring of hole images
 - Film had significant thickness variations due to rollers used in fabrication

